

### P7: Astronomy

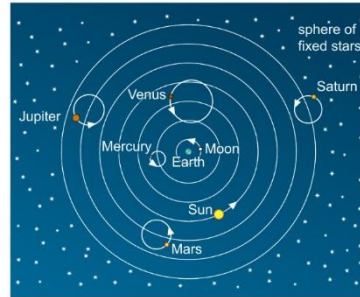
#### Lesson sequence

1. The solar system
2. Gravity and orbits
3. Life cycle of stars
4. Red-shift
5. Origin of the Universe

#### 1. The solar system

<b>Planet</b>	A celestial body moving in an elliptical orbit round a star.
<b>Natural Satellite</b>	Any celestial body in space that orbits around a larger body.
<b>Elliptical orbit</b>	The revolving of one object around another in an oval-shaped path called an ellipse.
<b>Geocentric model</b>	(often exemplified specifically by the Ptolemaic system) is a superseded description of the Universe with Earth at the centre. Under the geocentric model, the Sun, Moon, stars, and planets all orbited Earth.
<b>Heliocentric model</b>	Is a superseded description of the Universe with the Earth and planets revolve around the Sun at the centre of the Solar System.
<b>Telescope</b>	An optical instrument that makes distant objects appear magnified by using an arrangement of lenses or curved mirrors and lenses
<b>Dwarf planet</b>	A celestial body orbiting a star that is massive enough to be rounded by its own gravity. The gravitational field of a dwarf planet is not strong enough to clear the neighbourhood
<b>Asteroids</b>	Comprised of rock and metal, and are smaller than planets.

<b>Comets</b>	Balls of ice and dust in orbit around the Sun. The orbits of comets are different from those of planets - they are elliptical. A comet's orbit takes it very close to the Sun and then far away again.
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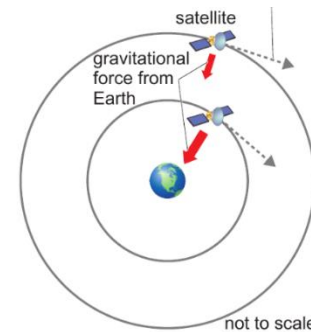
#### Exam-style question

Compare and contrast Ptolemy's and Copernicus' models of the Solar System. (4 marks)

#### 2. Gravity and orbits

<b>Weight</b>	A measure of the size of the pull of gravity on the object. Measured in Newtons (N)
<b>Mass</b>	A measure of how much matter there is in an object. Measured in Kilograms (Kg)
<b>Calculating Weight</b>	Weight = mass x gravitational field strength  Weight = Newtons (N) Mass = Kilograms (Kg) Gravitational field strength = N/Kg
<b>Gravitational field strength</b>	Is measured in newtons per kilogram (N/kg). The Earth's gravitational field strength is 9.8 N/kg. This means that for each kg of mass, an object will experience 9.8 N of force.

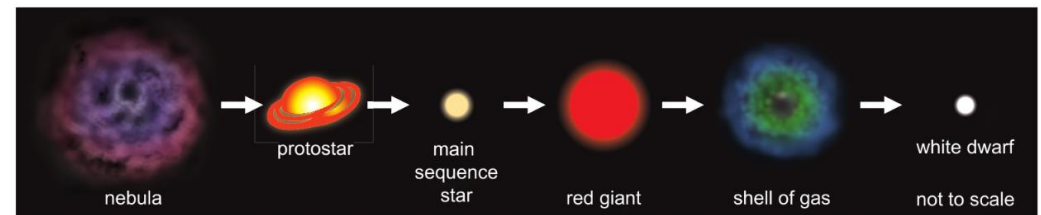
<b>Artificial Satellites</b>	A man-made body placed in orbit round the earth or another planet in order to collect information about it or for communication purposes.
<b>Velocity</b>	Its speed in a particular direction.
<b>Vector Quantity</b>	A vector describes a movement from one point to another. A vector quantity has both direction and magnitude (size).

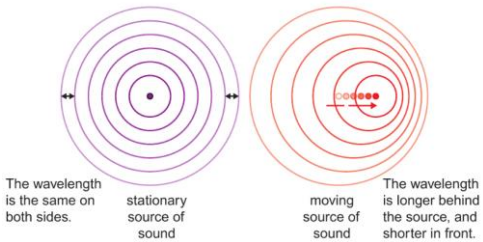


#### 3. Life cycle of stars

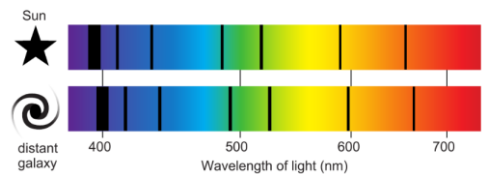
<b>Nebula</b>	A massive cloud of dust and gas in space
<b>Protostar</b>	A contracting mass of gas which represents an early stage in the formation of a star, before nucleosynthesis has begun.

<b>Fusion reactions</b>	A process where two or more nuclei combine to form an element with a higher atomic number (more protons in the nucleus). Fusion releases energy
<b>Electromagnetic radiation</b>	A wave of the electromagnetic field, propagating (radiating) through space, carrying electromagnetic radiation energy.
<b>Main sequence</b>	Most of the stars in the Universe are in the main sequence stage of their lives, a point in their stellar evolution where they're converting hydrogen into helium in their cores and releasing a tremendous amount of energy.
<b>Red giant</b>	A very large star of high luminosity and low surface temperature. Red giants are thought to be in a late stage of evolution when no hydrogen remains in the core to fuel nuclear fusion.
<b>White dwarf</b>	Is formed when a low-mass star has exhausted all its central nuclear fuel and lost its outer layers as a planetary nebula.



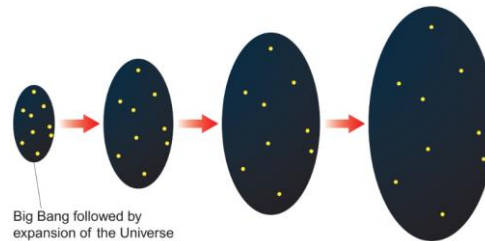


<b>Red supergiant's</b>	A red supergiant is an aging giant star that has consumed its core's supply of hydrogen fuel.
<b>Supernova</b>	A star that suddenly increases greatly in brightness because of a catastrophic explosion that ejects most of its mass.
<b>Black hole</b>	A region of space having a gravitational field so intense that no matter or radiation can escape.
<b>Neutron star</b>	A celestial object consisting of an extremely dense mass of neutrons, formed at the core of a supernova, where electrons and nuclei are compressed together so intensely by the force of gravity that protons and electrons merge together into neutrons.



#### 4. Red-shift

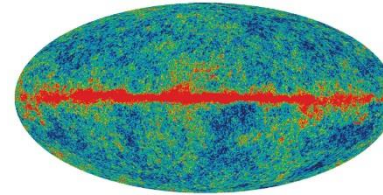
<b>Doppler effect</b>	An increase (or decrease) in the frequency of sound, light, or other waves as the source and observer move towards (or away from) each other.
<b>Pitch</b>	The quality of a sound governed by the rate of vibrations producing it;
<b>Red shift</b>	the displacement of spectral lines towards longer wavelengths (the red end of the spectrum) in radiation from distant galaxies and celestial objects.
<b>Red-shifted</b>	the wavelength of the light is stretched, so the light is seen as 'shifted' towards the red part of the spectrum.
<b>Blue-shift</b>	the displacement of the spectrum to shorter wavelengths in the light coming from distant celestial objects moving towards the observer.
<b>Universe</b>	all existing matter and space considered as a whole; the cosmos.



#### 5. Origin of the Universe

<b>Big bang theory</b>	It is the idea that the whole Universe and all the matter in it started out as a tiny point of concentrated energy about 13.5 billion years ago. The universe expanded from this point and is still expanding.
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<b>Steady state theory</b>	States that the Universe has always existed, and that the Universe is expanding and constantly creating matter as the Universe expands.
<b>Cosmic microwave background (CMB) radiation</b>	This comes from all directions in space and has a temperature of about $-270^{\circ}\text{C}$ . The CMBR is the remains of the thermal energy from the Big Bang, spread thinly across the whole Universe.



#### Exam-style question

Compare and contrast the Big Bang and Steady State theories. (3 marks)